

WATER-HAMMER JETTING EMANATING FROM IMPACTING DROPS

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ABSTRACT

This work studies the formation and evolution of focused jets. In particular, the dependence of jet formation and evolution on selected physical parameters is investigated. The data are obtained using the conservative sharp-interface method for the setups described in Castillo-Orozco *et al.* (2015).

INTRODUCTION

Liquid jets can be observed in a number of natural phenomena. A type of the liquid jets which is of a distinct interest due to its frequent usage in science and industry is a focused liquid jet. To optimally employ focused jets for specific use cases, the parametric dependencies of their formation and evolution must be well understood. A clear understanding of these phenomena requires insight into the physics of jet formation.

Experiments are limited in their capacity as they are often sensitive to the experimental design in an unexpected manner George (1990) and are never free of aleatoric uncertainties. The aforementioned drawbacks of physical experiments lead to the necessity to perform numerical simulations and to thereby specifically model the physics accurately.

However, the study of shear-flow configuration of the highly focused liquid jets with numerical simulations is also challenging. The core - the highly focused jet - is characterized by fine-scaled, quickly evolving structures. The periphery is characterized by a free surface and a bulk liquid, both carrying and refocusing the wave-dynamics such to form a highly focused liquid jet.

These challenges can be tackled with a very-low dissipative gradient-capturing flow solver, proper physical models of the liquid and gaseous phases, a sharp-interface representation and sufficient resolution.

However, when seeking to understand the parametric dependency of the jet formation and evolution, the actual challenge is to conduct a study campaign computationally efficiently. Therefore, computational resources need to be focused without loss of accuracy whenever possible and the axisymmetry of the configuration needs to be exploited. The approach has been proven to be qualitatively and quantitatively accurate for similar applications in a number of jet-formation stud-

ies, i.e., Zahoor *et al.* (2021), Saade *et al.* (2021), Peters *et al.* (2013) and Kyriazis *et al.* (2019).

For these reasons, the simulations are conducted using the level-set based sharp-interface multiresolution solver ALPACA Hoppe *et al.* (2022). This solver demonstrates its suitability to predict the multiphase-flow behavior in diverse scenarios, e.g., Kaiser *et al.* (2020); Hoppe *et al.* (2019); Fleischmann *et al.* (2019); Bußmann *et al.* (2023); Winter *et al.* (2023). The computational efficiency of the method demonstrated in Hoppe *et al.* (2022) allows to adequately resolve fine scales where essential for the representation of focused liquid jets.

In this work the solver's capability in exploiting the axisymmetry of the shear-flow configuration shall be explored. Note that the challenge therein lies in the fact that the highly-focused jet forms and evolves at the symmetry and not at the outer boundary, posing a challenge in terms of accuracy and robustness on a numerical solver.

The work is structured as follows: the physical model, numerical implementation, and numerical setup are described, followed by presentation and discussion of simulation results from a series of parametric studies.

PHYSICAL MODEL AND ITS NUMERICAL IMPLEMENTATION

Physical model

The compressible Navier-Stokes equations, including viscous, surface-tension effects and gravitational acceleration, hold:

$$\frac{\partial \mathbf{U}}{\partial t} + \nabla \mathbf{F}(\mathbf{U}) = \mathbf{S}, \quad (1)$$

where $\mathbf{U} = (\rho, \rho u_1, \rho u_2, \rho u_3, \rho e_t)^T$ is the vector of conservatives, $\mathbf{F} = (\rho \mathbf{u}, \rho \mathbf{u} \otimes \mathbf{u} - \mathbf{\Pi}, \rho \mathbf{e}_t \mathbf{u} - \mathbf{u} \cdot \mathbf{\Pi})^T$ is the vector of fluxes and $\mathbf{S} = (0, \rho \mathbf{g}, \rho \mathbf{g}^T \cdot \mathbf{u})$ is the vector of source terms. In this case, ρ is the non-dimensional density, $\mathbf{u}$ is the non-dimensional velocity, $\mathbf{\Pi}$ is the non-dimensional stress tensor, e_t is the non-dimensional specific total energy, p is the static,

non-dimensional pressure. The non-dimensionalization is described in detail in Hoppe *et al.* (2022).

The set of equations is closed with the Tait equation of state for the liquid phase:

$$p = p(\rho) = p_1 \left[\left(\frac{\rho}{\rho_0} \right)^\omega - 1 \right] + p_0, \quad (2)$$

where ω is the heat capacity ratio, ρ_0 is the reference density, p_0 is the reference pressure.

For the gaseous phase, the ideal-gas equation of state holds:

$$p = (\omega - 1)\rho e \quad (3)$$

Numerical implementation

A finite-volume approach on a Cartesian grid is applied to solve Eq. 1.

The two phases and the interface are treated by a sharp-interface method, see e.g. Schraner (2017); Schraner & Adams (2016); Schraner *et al.* (2016). For the time integration, the 3rd-order Runge-Kutta TVD scheme Gottlieb & Shu (1998) is utilized. An adaptive multi-resolution approach Harten (1994) is used to increase the computational efficiency.

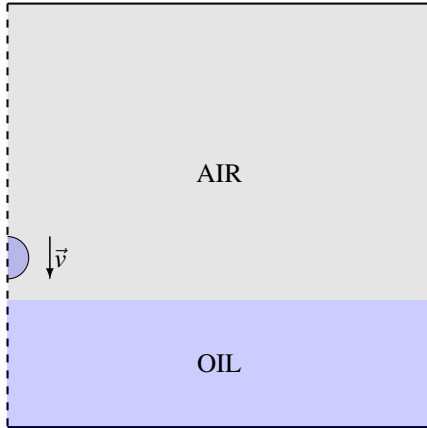


Figure 1. Initial setup.

Numerical setup

The numerical setup is designed to replicate the initial conditions of experiments from Castillo-Orozco *et al.* (2015). A rectangular domain is filled with air and silicone oil, see Fig. 1, so that a small drop of silicone oil surrounded by air is falling into a pool of silicone oil. The symmetry boundary condition is applied at the left domain boundary. The right and bottom boundaries are solid walls, and the top boundary uses a zero-gradient boundary condition. The simulation is run axisymmetrically. The gravitational acceleration is directed downwards.

The air and oil, except for the drop, are initially static.

RESULTS AND DISCUSSION

Verification

To demonstrate that ALPACA adequately captures the physical behavior of the liquid jet in a weakly-compressible

framework and to justify the resolution chosen for the subsequent numerical experiments, the simulations using the earlier introduced setup have been performed with effective resolutions of 1024×1024 , 2048×2048 , and 4096×4096 cells on the finest level. Fig. 2 depicts the shift of the jet tip as a function of time for the three simulations. While the results for the coarsest resolution diverges from the other time-evolutions shortly after $t = 0.02$, the trends of h for the two finest resolutions remain nearly identical until the jet reaches its peak. This justifies the choice of 2048×2048 cells for the subsequent simulations, as the maximum jet tip height is used as the key variable for the sensitivity analysis in the following sections.

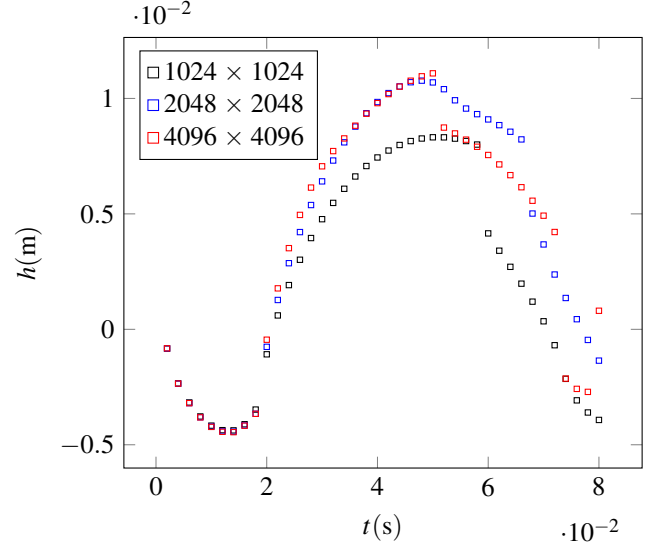


Figure 2. Jet-tip height h as a function of time at different effective resolutions

Effect of the surface tension

The effect of the following independent quantities is studied: surface tension, viscosity, initial velocity of the drop, and gravitational acceleration. This choice is based on the significant influence each of quantity has on jet behavior: surface tension and viscosity represent fluid properties, initial velocity characterizes the fluid state at the initial moment, and gravitational acceleration characterizes the external conditions.

This subsection focuses on surface tension. Figures 3 and 4 show the jet-tip height as a function of time, with results grouped by parameter value in ascending order for clarity (this grouping convention is used throughout this and the following subsections). From Figure 3, it is visible that the surface-tension coefficient starts to play a noticeable role in the jet evolution starting from the values around $\sigma = 10$ mN/m. For $\sigma = 10$ mN/m and $\sigma = 15$ mN/m, the formation of secondary droplets is observed, which are represented by the discontinuities of the function $h = h(t)$. However, with further increase of the surface tension coefficient, the tendency for the secondary droplet formation declines, as clearly visible in Figure 4.

Effect of the viscosity

Dynamic viscosity is another liquid parameter with a strong impact on jet behavior. As expected, the fastest jet formation occurs at the lowest viscosities (see Figure 5). For

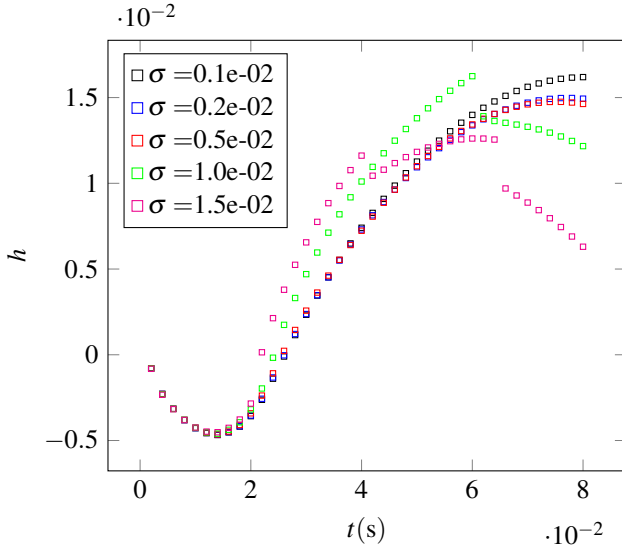


Figure 3. $h = h(t)$ for different surface-tension coefficients (lower values)

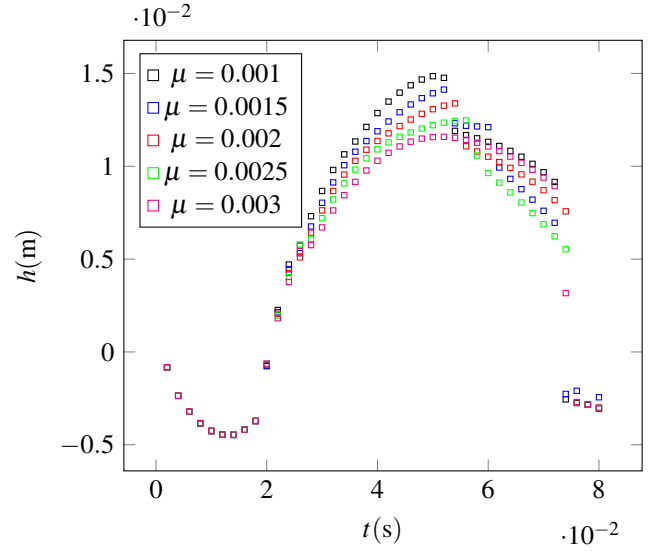


Figure 5. $h = h(t)$ for different dynamic viscosities (lower values)

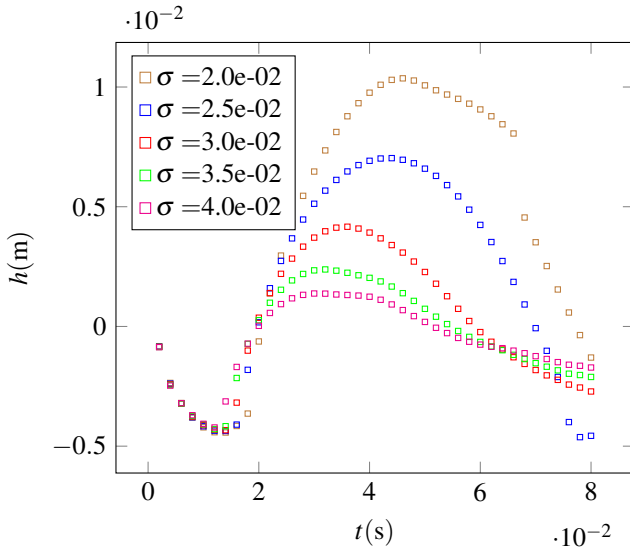


Figure 4. $h = h(t)$ for different surface-tension coefficient (higher values)

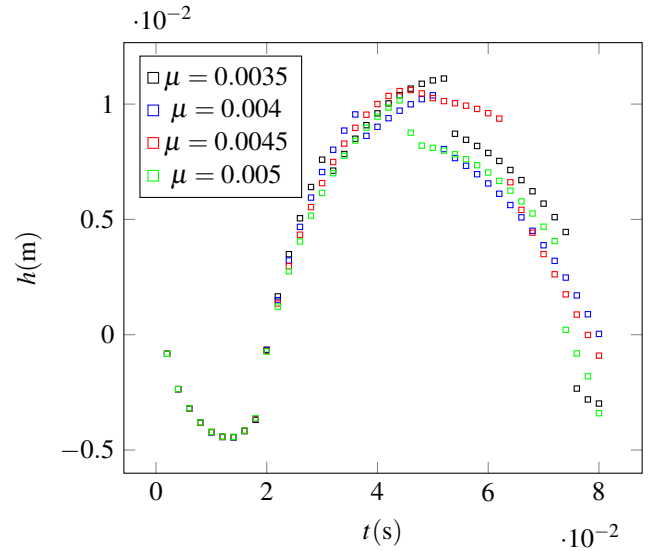


Figure 6. $h = h(t)$ for different dynamic viscosities (middle values)

most simulations in this subsection, the behavior follows this sequence: the jet forms and reaches its maximum height, the first secondary droplet forms, the main jet descends, and in some cases, a second secondary droplet forms (see the results for $\mu = 0.0035$ kg/ms, Figure 6, as an example). However, for $\mu = 0.003$ kg/ms, the secondary droplet does not form at the maximum height but only near the end of the simulation. The evolution profiles are nearly identical up to $t = 20$ ms. Afterwards jets with the lowest viscosities exhibit the most rapid growth to peak height. The secondary droplets form for all viscosities, earliest at $t \approx 40$ ms ($\mu = 0.0055$ kg/ms, Figure 7), latest at $t \approx 80$ ms ($\mu = 0.003$ kg/ms, Figure 5).

Effect of the initial velocity

Initial velocity, as a characteristic of the drop's initial-state, has a major impact on jet behavior. This impact is especially pronounced at lower velocities, as shown in Figure 8. At the lowest velocity $v = 1.0$ m/s, the jet development is

not characterized by secondary droplet formation. However, at higher velocities, the maximum height increases and one or two consecutive secondary droplets typically form. Notably, for velocities above $v = 3.0$ m/s (Figure 9), further velocity increases have less influence on jet behavior compared to the pronounced changes observed for similar velocity increments at lower velocities.

Effect of the gravitational acceleration

To examine the influence of external conditions on jet formation, simulations with different gravitational accelerations have been conducted. Similar to the sensitivity to initial velocity observed in the initial-velocity parametric study, early jet development shows high sensitivity to gravity until reaching maximum height (Figure 10) but is much less sensitive at higher values (Figure 11). Moreover, at the largest values, a second jet begins to form after the initial jet subsides, with the effect being especially strong for $g = 45.0$ m/s².

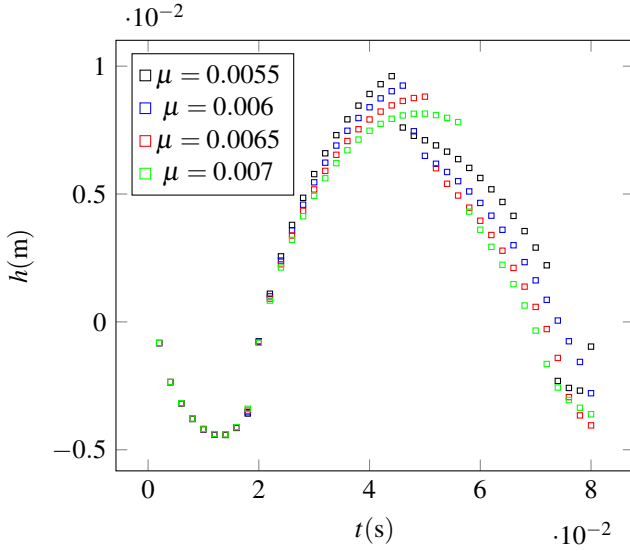


Figure 7. $h = h(t)$ for different dynamic viscosities (higher values)

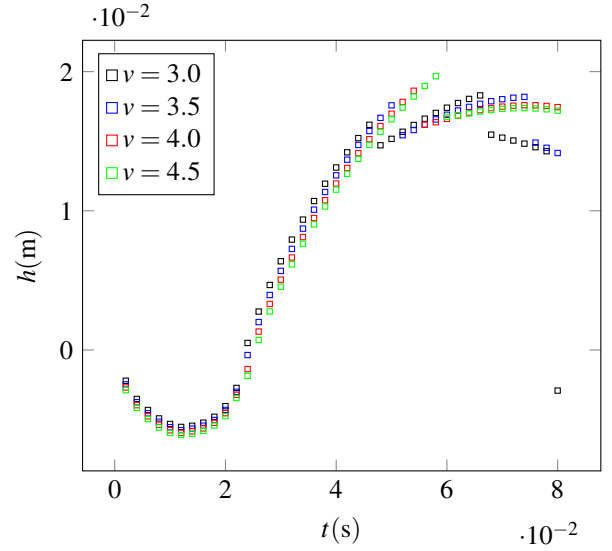


Figure 9. $h = h(t)$ for different velocities (higher values)

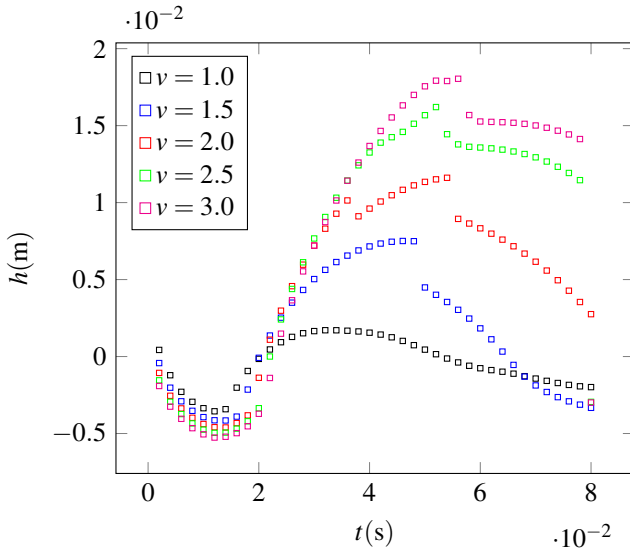


Figure 8. $h = h(t)$ for different velocities (lower values)

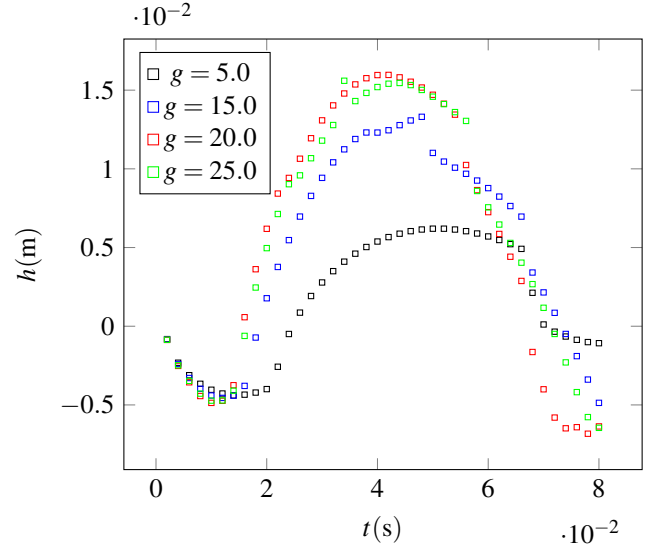


Figure 10. $h = h(t)$ for different gravitational accelerations (lower values)

DISCUSSION AND CONCLUSION

The following evolution stages are typically observed for the jets investigated in this work: (1) impact cavity formation; (2) jet growth to maximum height; (3) secondary droplet formation; (4) jet decline sometimes accompanied by additional secondary droplet formation. Our results for jet evolution are qualitatively consistent with the ones obtained by Castillo-Orozco *et al.* (2015). This demonstrates the capability of the numerical simulations obtained using the conservative sharp-interface method to adequately represent the behavior of a drop-triggered liquid jet, including the formation of the secondary drops.

The parametric studies presented in this work capture the dependence of the jet-profile evolution on particular characteristics of the fluids, the ambient conditions and the initial setup. The evolution characteristics affected involve the speed of the jet formation, the moments of reaching the peak and drop formation and whether the secondary drops are formed at all. Quantifying the influence of individual parameters on specific evolution characteristics remains a subject for future

investigation.

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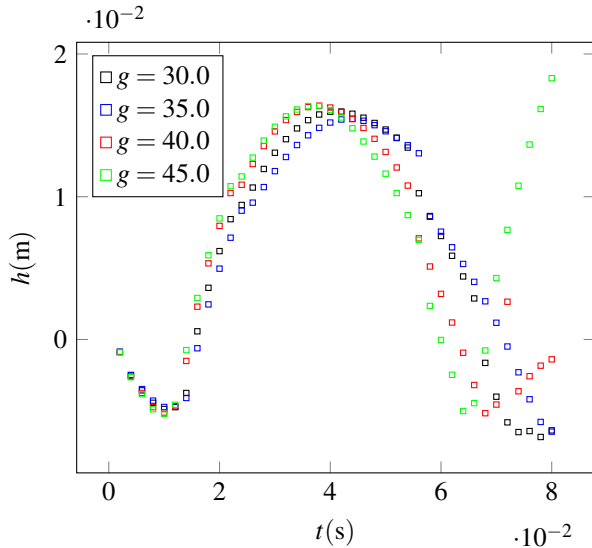


Figure 11. $h = h(t)$ for different gravitational accelerations (higher values)

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